Tethys Scatterometry Rev 21

R. West

December 13, 2005

• Sequence: s18

• Rev: 021

Observation Id: te_021_1

• Target Body: Tethys

1 Introduction

This memo describes one of the Cassini RADAR activities for the s18 sequence of the Saturn Tour. A sequence design memo provides the science context of the scheduled observations, an overview of the pointing design, and guidlines for preparing the RADAR IEB.

This IEB is for the second Tethys distant scatterometer observation around Feb 25 2006. The usual warmup parameters are used during the first three hours as shown in table 3

2 CIMS and Division Summary

Each RADAR observation is represented to the project by a set of requests in the Cassini Information Management System (CIMS). The CIMS database contains requests for pointing control, time, and data volume. The CIMS requests show a high-level view of the sequence design.

The CIMS requests form the basis of a pointing design built using the project pointing design tool (PDT). The details of the pointing design are shown by the PDT plots on the corresponding tour sequence web page. (See https://cassini.jpl.nasa.gov/radar.) The RADAR pointing sequence is ultimately combined with pointing sequences from other instruments to make a large merged c-kernel. C-kernels are files containing spacecraft attitude data.

A RADAR tool called RADAR Mapping and Sequencing Software (RMSS) reads the merged c-kernel along with other navigation data files, and uses these data to produce a set of instructions for the RADAR observation. The RADAR instructions are called an Instrument Execution Block (IEB). The IEB is produced by running RMSS with a radar config file that controls the process of generating IEB instructions for different segments of time. These segments of time are called divisions with a particular behavior defined by a set of division keywords in the config file. Table 1 shows a summary of the divisions used in this observation. Subsequent sections will show and discuss the keyword selections made for each division. Each division table shows a set of nominal parameters that are determined by the operating mode (eg., distant scatterometry, SAR low-res inbound). The actual division parameters from the config file are also shown, and any meaningful mismatches are flagged.

Division	Name	Start	Duration	Data Vol	Comments	
a	distant_warmup	-2:56:0.0	00:08:0.0	0.1	Warmup	
b	distant_warmup	-2:48:0.0	02:50:0.0	10.1	Blind Warmup	
С	scat_compressed	00:02:0.0	00:10:0.0	4.2	Scatterometer off-target receive only compressed	
d	distant_radiometer	00:12:0.0	00:06:0.0	0.4	Radiometer during final turn transition	
e	scat_compressed	00:18:0.0	00:02:0.0	0.8	Scatterometer on-target receive only compressed	
f	distant_scatterometer	00:20:0.0	00:23:0.0	282.9	Scatterometer target-center stare with tone	
g	scat_compressed	00:43:0.0	00:10:0.0	4.2	Scatterometer on-target receive only compressed	
h	distant_radiometer	00:53:0.0	00:18:0.0	1.1	Radiometer during final stare and turn to waypoint	
Total				303.8		

Table 1: Division summary. Data volumes (Mbits) are estimated from maximum data rate and division duration.

Div	Alt (km)	Slant range (km)	B3 Size (target dia)	B3 Dop. Spread (Hz)
a	158966	off target	0.97	off target
b	161134	off target	0.98	off target
С	226165	off target	1.38	off target
d	231013	off target	1.41	off target
e	233970	233970	1.43	478
f	234964	234964	1.43	491
g	246680	246680	1.51	639
h	251935	251935	1.54	699

Table 2: Division geometry summary. Values are computed at the start of each division. B3 Doppler spread is for two-way 3-dB pattern. B3 size is the one-way 3-dB beamwidth

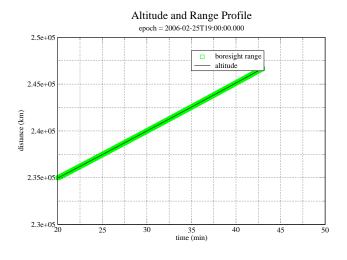


Figure 1: Div F: Altitude and range to the boresight point

Name	Nominal	Actual	Mismatch	Comments
mode	radiometer	radiometer	no	
start_time (min)	varies	-176.0	no	
end_time (min)	varies	-168.0	no	
time_step (s)	varies	1800.0	no	Used by radiome-
				ter only modes -
				saves commands
bem	00100	00100	no	
baq	don't care	5	no	
csr	6	6	no	6 - Radiometer
				Only Mode
noise_bit_setting	don't care	4.0	no	
dutycycle	don't care	0.38	no	
prf (KHz)	don't care	1.00	no	
number_of_pulses	don't care	8	no	
n_bursts_in_flight	don't care	1	no	
percent_of_BW	don't care	100.0	no	
auto_rad	on	on	no	
rip (ms)	34.0	34.0	no	
max_data_rate	0.250	0.248	yes	Kbps - set for
				slowest burst pe-
				riod
interleave_flag	off	off	no	
interleave_duration (min)	don't care	10.0	no	

Table 3: te_021_1 div_a distant_warmup block

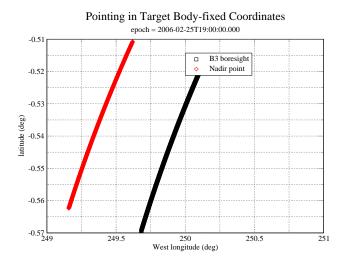


Figure 2: Div's F: Stare in target body-fixed coordinates

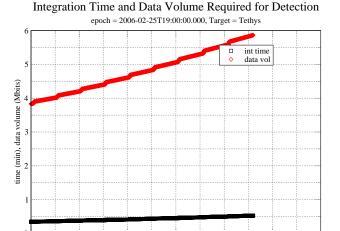


Figure 3: Scatterometry Div F: Detection integration time required for a single point detection using optimal chirp bandwidth

time (min)

45

3 Div's C-H: Tethys Radiometry/Scatterometry

Tethys is 530 km in radius, Allocated data volume is 290 Mbits. Figures 1 and 2 show the pointing design for the scatterometry stare from the merged ckernel. The angular size of the target is about 4.5 mrad during this division. The beam 3 beamwidth is 6 mrad and a single point stare is used.

The division parameters for the compressed scatterometer receive only integrations are shown in table 4 and the target integration division is shown in table 5.

3.1 Scatterometer Receive Only Measurements

Before the actual scatterometer integration, two measurements will be collected in receive only mode. One is collected before the spacecraft slews on target to provide a cold space reference level. The second is collected while pointing at the target. Both will last 10 minutes and use the compressed mode to acquire a solid noise measurement without using much data volume. These measurements can then be used to determine system gain and receiver noise temperature for the scatterometer mode right before the actual target measurement. The division PRF and number of pulses (1202 Hz and 150 respectively) are chosen to fill the science data buffer. These parameters give the best performance possible from the compressed mode.

3.2 Scatterometer Performance

The detection performance is shown in figures 3, 4, and 5. Figure 5 shows that range processing will likely be difficult due to range ambiguities. On the other hand, disk integrated results should be very stable.

The maximum doppler spread in Div f is 639 Hz which comes from rotation and spacecraft motion. The PRF needs to be higher than the doppler spread to support potential range-doppler processing, and is set by division parameter to 977 Hz. With this PRF, the range amiguity spacing is 153 km while Tethys is 530 km in radius. The range-spread of the beam depends on where it is pointed. For target centered pointing the cosine law can be applied to solve the geometry. At 234964 km range, the range-spread is 529 km.

Unfortunately, this is still larger than the range ambiguity spacing, so it will not be possible to collect unambiguous data. Due to the range ambiguity problem, a chirp transmission will not be used for this observation. Instead, a tone transmission will be used in all of the scatterometer divisions. Even though a tone transmission is used, the PRF is still set higher than the doppler spread to separate out the grating pattern due to the pulse structure of each burst transmission. One might consider doing ranging to obtain tighter incidence angle results than the doppler processing can achieve, however, the signal variance is too high at the required resolution (see Fig. 5).

Name	Nominal	e	g	Mismatch	Comments
mode	scat_compressed	scatterometer	scatterometer	yes	
start_time (min)	varies	18.0	43.0	no	
end_time (min)	varies	20.0	53.0	no	
time_step (s)	don't care	6.0	2.0	no	Set by valid time calculation
bem	00100	00100	00100	no	
baq	3	3	3	no	3 - PRI summation
csr	1	1	1	no	0 - normal operation with fixed attenuator set to match Phoebe for easier cross-calibration
noise_bit_setting	4.0	4.0	4.0	no	9 dB setting used by all low SNR scatterometry
dutycycle	0.70	0.70	0.70	no	
prf (KHz)	1.20	1.20	1.20	yes	Set with num pulses to fill science data buffer
number_of_pulses	150	150	150	no	Set with the PRF to fill the science data buffer - Only 2 PRI's worth of data are downlinked.
n_bursts_in_flight	1	1	1	no	
percent_of_BW	100.0	0.0	0.0	yes	
auto_rad	on	on	on	no	
rip (ms)	34.0	34.0	34.0	no	
max_data_rate	8.000	8.000	8.000	no	
interleave_flag	off	off	off	no	
interleave_duration (min)	don't care	10.0	10.0	no	

Table 4: te_021_1 div_eg scat_compressed block

Name	Nominal	Actual	Mismatch	Comments
mode	scatterometer	scatterometer	no	
start_time (min)	varies	20.0	no	
end_time (min)	varies	43.0	no	
time_step (s)	don't care	12.0	no	Used when BIF > 1, otherwise set by valid time cal- culation
bem	00100	00100	no	
baq	5	5	no	
csr	0	0	no	0 - normal operation with fixed attenuator set to match Phoebe for easier cross-calibration
noise_bit_setting	4.0	4.0	no	Scat signal set higher than ALT/SAR
dutycycle	0.70	0.70	no	
prf (KHz)	varies	0.98	no	Set to cover doppler spread and to allow CSF = integer multiple
number_of_pulses	varies	60	no	depends on PRF choice (can have more shorter pulses)
n_bursts_in_flight	varies	3	no	Used to increase PRF and data rate at long range
percent_of_BW	0.0	0.0	no	
auto_rad	on	on	no	
rip (ms)	34.0	34.0	no	
max_data_rate	200.000	205.000	yes	Kbps - determines burst period
interleave_flag	off	off	no	
interleave_duration (min)	don't care	10.0	no	

Table 5: te_021_1 div_f distant_scatterometer block

Figure 4: Outbound observation Div F: Normalized estimated signal standard deviation for a disk integrated observation using optimal chirp bandwidth and assuming all the bursts occur at minimum range, and 15 minutes away from minimum range.

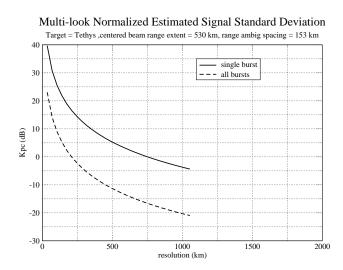


Figure 5: Outbound observation Div F: Normalized estimated signal standard deviation for a range/doppler cell as a function of resolution. Range/doppler resolution elements are both set equal to the specified resolution. Results are shown for a single burst, and for all the bursts in this division. Calculations are performed using the geometry at the start of the division. The presence of ambiguities are not shown.

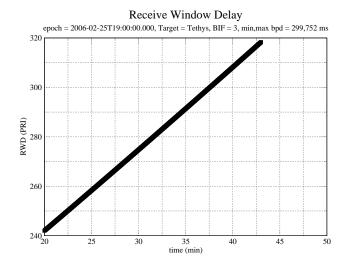


Figure 6: Div F: Inbound scatterometer receive window delay. Subtitle shows the minimum and maximum burst periods that are in principle compatible with the division selected number of bursts in flight.

4 Div H: Tethys Radiometry

Following the scatterometer stare, the radar beam moves off target to the next waypoint. There is not enough time for any radiometry raster scans, so the scatterometry stare will supply the radiometry on target observation. This data along with the cold sky data and the internal reference load data will be used to calibrate the radiometer. The radiometer calibration also contributes to the scatterometer calibration. Division parameters for the radiometry raster are shown in table 5

5 Revision History

1. Dec 13, 2005: Initial Release

Name	Nominal	Actual	Mismatch	Comments
mode	radiometer	radiometer	no	
start_time (min)	varies	53.0	no	
end_time (min)	varies	71.0	no	
time_step (s)	varies	1800.0	no	Used by radiome-
				ter only modes
bem	00100	00100	no	
baq	don't care	5	no	
csr	6	6	no	
noise_bit_setting	don't care	4.0	no	
dutycycle	don't care	0.38	no	
prf (KHz)	don't care	1.00	no	
number_of_pulses	don't care	8	no	
n_bursts_in_flight	don't care	1	no	
percent_of_BW	don't care	100.0	no	
auto_rad	on	on	no	
rip (ms)	34.0	34.0	no	starting value for
max_data_rate	1.000	0.992	yes	auto-rad 1 Kbps - 1 s burst
max_uata_uate	1.000	0.992	yes	period which is
				adequate for slow
				radiometer scans
interleave_flag	off	off	no	
interleave_duration (min)	don't care	10.0	no	

Table 6: te_021_1 div_h distant_radiometer block

6 Acronym List

AL	Acronym List

ALT Altimeter - one of the radar operating modes

BAQ Block Adaptive Quantizer

CIMS Cassini Information Management System - a database of observations

Ckernel NAIF kernel file containing attitude data

DLAP Desired Look Angle Profile - spacecraft pointing profile designed for optimal SAR performance

ESS Energy Storage System - capacitor bank used by RADAR to store transmit energy

IEB Instrument Execution Block - instructions for the instrument

ISS Imaging Science Subsystem

IVD Inertial Vector Description - attitude vector data

IVP Inertial Vector Propagator - spacecraft software, part of attitude control system

INMS Inertial Neutral Mass Spectrometer - one of the instruments

NAIF Navigation and Ancillary Information Facility

ORS Optical Remote Sensing instruments

PDT Pointing Design Tool
PRI Pulse Repetition Interval
PRF Pulse Repetition Frequency

RMSS Radar Mapping Sequencing Software - produces radar IEB's

SAR Synthetic Aperture Radar - radar imaging mode

SNR Signal to Noise Ratio

SOP Science Operations Plan - detailed sequence design

SOPUD Science Operations Plan Update - phase of sequencing when SOP is updated prior to actual sequencing

SSG SubSequence Generation - spacecraft/instrument commands are produced

SPICE Spacecraft, Instrument, C-kernel handling software - supplied by NAIF to use NAIF kernel files.

TRO Transmit Receive Offset - round trip delay time in units of PRI